

# Small Extra Dimensions and Enhanced Symmetries in Orbifold Compactifications

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`arXiv:0803.4501`

`arXiv:0810.xxxx`

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- Higher dimensional setups provide promising frameworks for supersymmetric GUTs
- Size of the extra dimensions undetermined  
⇒ Moduli problem
- Casimir energy induces a nontrivial potential
- Compactification often leads to additional  $U(1)$  symmetries
- Combination of Casimir energy with gauge symmetry breaking can lead to small extra dimensions

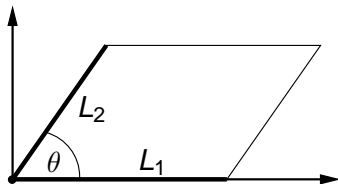
- 1 Example: An Orbifold GUT Model
- 2 Casimir Energy
- 3 Stabilisation

# Orbifold Compactification

- **Starting point:** higher-dimensional setup

**Here:** two extra dimensions, compactified on a torus

- Torus specified by the volume  $\mathcal{A}$  and shape  $\tau$



$$\mathcal{A} = L_1 L_2 \sin \theta$$

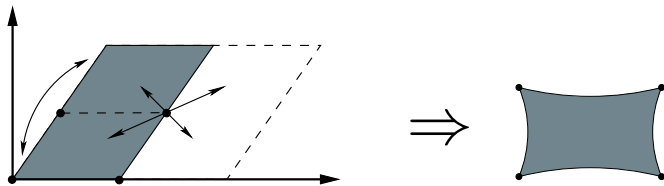
$$\tau = L_2 / L_1 e^{i\theta}$$

# Orbifold Compactification

- **Starting point:** higher-dimensional setup

**Here:** two extra dimensions, compactified on a torus

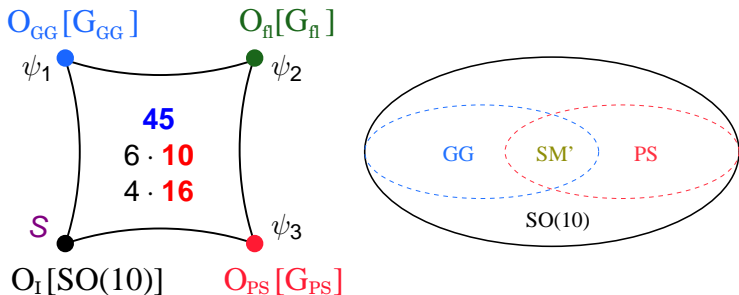
- Torus specified by the volume  $\mathcal{A}$  and shape  $\tau$



- Impose symmetry  $\mathbb{Z}_2 : y \rightarrow -y$
- Values for  $\tau, \mathcal{A}$ ?
- Casimir energy of bulk fields induces nontrivial potential
- Supersymmetry  $\Rightarrow$  vanishing Casimir energy  $\Rightarrow$  SUSY breaking

# Gaugino Mediation in a 6D Orbifold GUT Model

Asaka, Buchmüller, Covi, Phys. Lett. **B563** (2003)



## Gaugino Mediation

Kaplan, Kribs, Schmaltz, Phys. Rev. **D62** (2000)

Chacko, Luty, Nelson, Ponton, JHEP **01** (2000)

In general: soft masses for all bulk fields

Gaugino masses:  $m_g = \frac{\lambda\mu}{\Lambda^2\mathcal{A}}$       Scalar masses:  $m_H^2 = -\frac{\lambda'\mu^2}{\Lambda^2\mathcal{A}}$

# Casimir Energy

- Consider one-loop Casimir energy of a real scalar field
- Geometry:  $T^2/\mathbb{Z}_2$
- Fields can be either even or odd wrt a  $\mathbb{Z}_2$  symmetry
- Only fields which couple to SUSY breaking brane contribute
- **Boundary conditions** encoded in  $\alpha, \beta \in \{0, 1/2\}$

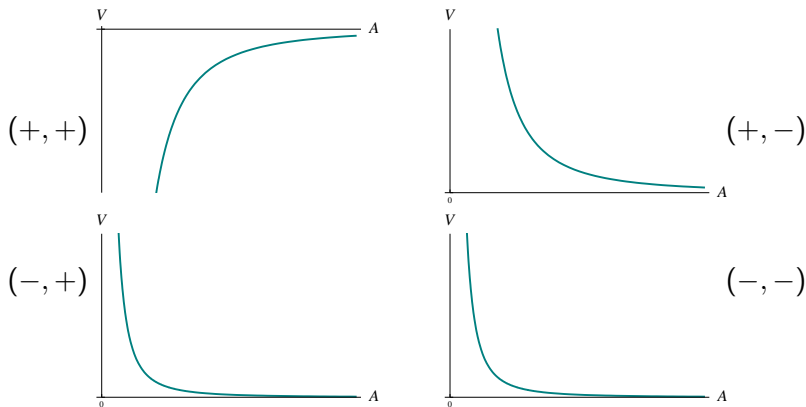
⇒ Four different contributions,

$$V_M^{\alpha, \beta} = \frac{1}{2} \left[ \sum \right]_{m, n}^{(\alpha, \beta)} \int \frac{d^4 k_E}{(2\pi)^4} \log \left( k_E^2 + \mathcal{M}_{m, n}^2 + M^2 \right)$$

$$\mathcal{M}_{m, n}^2 = \frac{4(2\pi)^2}{\mathcal{A}_{T^2}} |n + \beta - \tau(m + \alpha)|^2$$

- Zeta function regularisation

# Casimir Energy - Volume



- Sign and strength of Casimir force depends on boundary conditions
- General potential:  $a V^{(+,+)} + b V^{(+,-)} + c V^{(-,+)} + d V^{(-,-)}$



# Casimir Energy

- Analytical behaviour for small volume with  $\tau_1$  and  $\tau_2$  in the minimum:

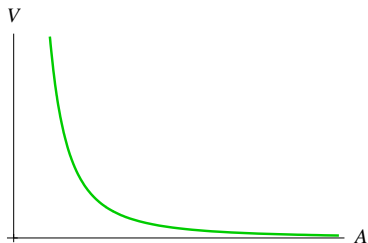
$$V_M^{(0,0)}(\tau_1 = \frac{1}{2}, \tau_2 = \frac{1}{2}, \mathcal{A}) \simeq -\frac{4\pi^3}{945\mathcal{A}^2} + \frac{\pi M^2}{360\mathcal{A}} + \mathcal{O}(M^4)$$

- Contributions for bosons and fermions come with opposite sign  
⇒ Leading term cancels within supermultiplet
- $M^2 = M_{\text{SUSY}}^2 + m_{\text{soft}}^2$
- **Leading term in supermultiplet  $\propto m_{\text{soft}}^2$**

# Casimir Energy in the Orbifold GUT Model

- Resulting potential in considered model

$$V_H \simeq - \frac{\pi}{36} \frac{\mu^2 \lambda'}{\Lambda^2 \mathcal{A}^2}$$



⇒ Can achieve repulsive force at short distances  
But: Need additional ingredient for stabilisation

# Breaking of $U(1)_X$

- 4D gauge symmetry:  $G_{SM'} = SU(3)_c \otimes SU(2)_L \otimes U(1)_Y \otimes U(1)_X$
- Vev  $\langle \Phi \rangle$  breaks the additional  $U(1)_X$   
 $\Rightarrow$  Bulk mass  $M \sim g_6 \langle \Phi \rangle$
- Quantum corrections generically induce Fayet-Iliopoulos terms at the fixed points  
Lee, Nilles, Zucker, Nucl.Phys.B **680** (2004)  
Buchmüller, Lüdeling, Schmidt, JHEP **0709** (2007)
- Localised FI terms can induce vev for bulk fields in turn
- D-flatness implies  $\mathcal{A} \langle \Phi \rangle^2 \sim C \Lambda^2$ ,  $C \ll 1$

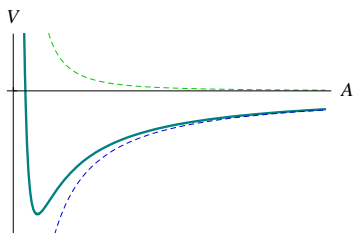
- **Classical** contribution to the vacuum energy density

$$\begin{aligned}V^{(0)} &= -\frac{\lambda''}{\Lambda^4} \int d^4\theta \langle S^\dagger S \Phi^\dagger \Phi \rangle \\ &= -\lambda'' \frac{\mu^2 C}{\mathcal{A}}\end{aligned}$$

- attractive for  $\lambda'' > 0$
- Combine with the repulsive Casimir energy

$$V_{\text{tot}} = V^{(0)} + V^{(1)} = -\frac{\pi}{36} \frac{\mu^2 \lambda'}{\Lambda^2 \mathcal{A}^2} - \frac{\lambda'' \mu^2 C}{\mathcal{A}}$$

# Volume Stabilisation

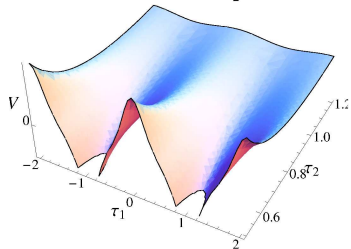
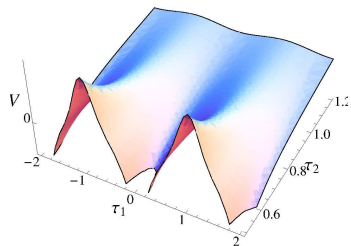
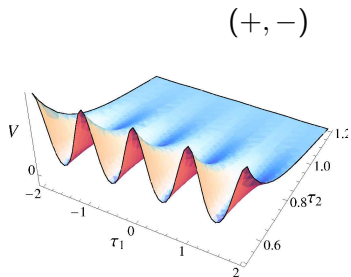
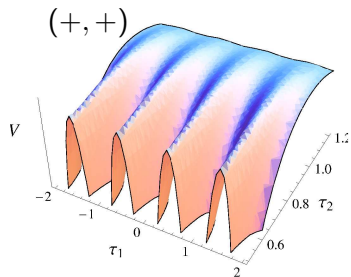


- Stable minimum at

$$\mathcal{A}_{\min} = -\frac{\pi\lambda'}{36\lambda''} \frac{1}{M^2} \lesssim \frac{1}{M^2}$$

- Independent of supersymmetry breaking scale  $\mu^2$
- Cosmological constant has to be tuned to zero by a brane cosmological term

# Casimir Energy - Shape

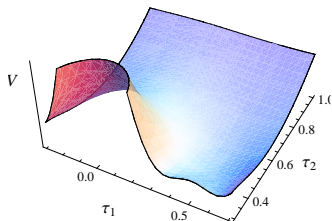


# Casimir Energy - Shape

- Casimir energy invariant under modular transformations

$$\tau \rightarrow \frac{a\tau + b}{c\tau + d}, \quad ad - bc = 1$$

- Modular transformation can have fixed points which correspond to extrema in the effective potential
- In our case we have  $c = 0 \pmod{2}$  and  $d = 1 \pmod{2}$
- These transformations have a fixed point at  $\tau_1 = \tau_2 = 1/2$



- Equivalent to  $R_1 = \sqrt{2}R_2$  and  $\theta = \pi/4 \Rightarrow$  Root lattice of  $SO(5)$

- Extra dimensions can be stabilised by interplay of gauge and supersymmetry breaking
- Compactification scale naturally of  $\mathcal{O}(M_{\text{GUT}})$  independently of supersymmetry breaking scale  $\mu^2$
- Leads to consistent picture of Orbifold GUTs
- Shape moduli stabilised at symmetry enhanced point



# Zeta function regularisation

$$V = - \left. \frac{d\zeta(s)}{ds} \right|_{s=0}$$

where

$$\zeta(s) = \frac{1}{2} \left[ \sum \right]_{m,n} \mu_r^{2s} \int \frac{d^4 k_E}{(2\pi)^4} \left( k_E^2 + \frac{4}{R_Z^2} [e^2(m + \alpha)^2 + (n + \beta)^2] + M^2 \right)^{-s}$$

# Casimir Energy

$$\begin{aligned}
 V_M^{\alpha,\beta} = & + \frac{M^6 \mathcal{A}}{3072\pi^3} \left[ \frac{11}{12} - \log\left(\frac{M}{\mu_r}\right) \right] \\
 & - \frac{M^4}{64\pi^2} \left[ \frac{3}{4} - \log\left(\frac{M}{\mu_r}\right) \right] \delta_{\alpha 0} \delta_{\beta 0} \\
 & - \frac{M^3 \tau_2^{3/2}}{4\pi^3 \mathcal{A}^{1/2}} \sum_{p=1}^{\infty} \frac{\cos(2\pi p \alpha)}{p^3} \mathcal{K}_3\left(p \frac{\sqrt{\mathcal{A}M}}{2\sqrt{\tau_2}}\right) \\
 & - \frac{32}{\mathcal{A}^2 \tau_2^2} \sum_{p=1}^{\infty} \sum_{m=0}^{\infty} \frac{1}{2^{\delta_{\alpha 0} \delta_{m 0}}} \frac{\cos(2\pi p(\beta - (m + \alpha)\tau_1))}{p^{5/2}} \left( \tau_2^2 (m + \alpha)^2 + \frac{\mathcal{A} \tau_2 M^2}{(4\pi)^2} \right)^{5/4} \\
 & \mathcal{K}_{5/2} \left( 2\pi p \sqrt{\tau_2^2 (m + \alpha)^2 + \frac{\mathcal{A} \tau_2 M^2}{(4\pi)^2}} \right)
 \end{aligned}$$

- Dependence on regularization scale  $\mu_r$  remnant of divergent bulk and brane cosmological terms